

e

$$J = \frac{1}{2}$$

e MASS (atomic mass units u)

The primary determination of an electron's mass comes from measuring the ratio of the mass to that of a nucleus, so that the result is obtained in u (atomic mass units). The conversion factor to MeV is more uncertain than the mass of the electron in u; indeed, the recent improvements in the mass determination are not evident when the result is given in MeV. In this datablock we give the result in u, and the following datablock in MeV.

VALUE (10^{-6} u)	DOCUMENT ID	TECN	COMMENT
548.5799110±0.0000012	MOHR 99	RVUE	1998 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
548.5799092±0.0000004	¹ BEIER 02	CNTR	Penning trap
548.5799111±0.0000012	² FARNHAM 95	CNTR	Penning trap
548.579903 ±0.000013	COHEN 87	RVUE	1986 CODATA value
¹ BEIER 02 compares Larmor frequency of the electron bound in a $^{12}\text{C}^{5+}$ ion with the cyclotron frequency of a single trapped $^{12}\text{C}^{5+}$ ion.			
² FARNHAM 95 compares cyclotron frequency of trapped electrons with that of a single trapped $^{12}\text{C}^{6+}$ ion.			

e MASS

The conversion from u (atomic mass units, see the above datablock) to MeV is 931.494013 ± 0.000037 MeV/u. The conversion error dominates the precision quoted in the following entries.

VALUE (MeV)	DOCUMENT ID	TECN	COMMENT
0.510998902±0.000000021	MOHR 99	RVUE	1998 CODATA value
• • • We do not use the following data for averages, fits, limits, etc. • • •			
0.510998901±0.000000020	^{3,4} BEIER 02	CNTR	Penning trap
0.510998903±0.000000020	^{3,5} FARNHAM 95	CNTR	Penning trap
0.510998895±0.000000024	³ COHEN 87	RVUE	1986 CODATA value
0.5110034 ±0.0000014	COHEN 73	RVUE	1973 CODATA value
³ Converted to MeV using the 1998 CODATA value of the conversion constant, 931.494013 ± 0.000037 MeV/u.			
⁴ BEIER 02 compares Larmor frequency of the electron bound in a $^{12}\text{C}^{5+}$ ion with the cyclotron frequency of a single trapped $^{12}\text{C}^{5+}$ ion.			
⁵ FARNHAM 95 compares cyclotron frequency of trapped electrons with that of a single trapped $^{12}\text{C}^{6+}$ ion.			

$(m_{e^+} - m_{e^-}) / m_{\text{average}}$

A test of *CPT* invariance.

VALUE	CL%	DOCUMENT ID	TECN	COMMENT
$<8 \times 10^{-9}$	90	⁶ FEE	93	CNTR Positronium spectroscopy
$<4 \times 10^{-8}$	90	CHU	84	CNTR Positronium spectroscopy

⁶ FEE 93 value is obtained under the assumption that the positronium Rydberg constant is exactly half the hydrogen one.

$|q_{e^+} + q_{e^-}|/e$

A test of *CPT* invariance. See also similar tests involving the proton.

VALUE	DOCUMENT ID	TECN	COMMENT
$<4 \times 10^{-8}$	⁷ HUGHES	92	RVUE
$<2 \times 10^{-18}$	⁸ SCHAEFER	95	THEO Vacuum polarization
$<1 \times 10^{-18}$	⁹ MUELLER	92	THEO Vacuum polarization

⁷ HUGHES 92 uses recent measurements of Rydberg-energy and cyclotron-frequency ratios.

⁸ SCHAEFER 95 removes model dependency of MUELLER 92.

⁹ MUELLER 92 argues that an inequality of the charge magnitudes would, through higher-order vacuum polarization, contribute to the net charge of atoms.

e MAGNETIC MOMENT ANOMALY

$\mu_e/\mu_B - 1 = (g-2)/2$

For the most accurate theoretical calculation, see KINOSHITA 81.

VALUE (units 10^{-6})	DOCUMENT ID	TECN	CHG	COMMENT
$1159.6521869 \pm 0.0000041$	10 MOHR	99	RVUE	1998 CODATA value
$<1159.6521869 \pm 0.0000041$	<10 COHEN	87	RVUE	1986 CODATA value
$1159.6521884 \pm 0.0000043$	VANDYCK	87	MRS	– Single electron
$1159.6521879 \pm 0.0000043$	VANDYCK	87	MRS	+ Single positron

¹⁰ The CODATA value assumes the $g/2$ values for e^+ and e^- are equal, as required by *CPT*.

$(g_{e^+} - g_{e^-}) / g_{\text{average}}$

A test of *CPT* invariance.

<u>VALUE</u> (units 10^{-12})	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
-0.5 ± 2.1		11 VANDYCK 87	MRS	Penning trap
• • • We do not use the following data for averages, fits, limits, etc. • • •				
< 12	95	12 VASSERMAN 87	CNTR	Assumes $m_{e^+} = m_{e^-}$
22 ± 64		SCHWINBERG 81	MRS	Penning trap
11 VANDYCK 87 measured $(g_-/g_+) - 1$ and we converted it.				
12 VASSERMAN 87 measured $(g_+ - g_-)/(g-2)$. We multiplied by $(g-2)/g = 1.2 \times 10^{-3}$.				

e ELECTRIC DIPOLE MOMENT

A nonzero value is forbidden by both *T* invariance and *P* invariance.

<u>VALUE</u> (10^{-26} ecm)	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
0.069 ± 0.074		REGAN 02	MRS	^{205}TI beams
• • • We do not use the following data for averages, fits, limits, etc. • • •				
0.18 ± 0.12 ± 0.10		13 COMMINS 94	MRS	^{205}TI beams
– 0.27 ± 0.83		13 ABDULLAH 90	MRS	^{205}TI beams
– 14 ± 24		CHO	NMR	TI F molecules
– 1.5 ± 5.5 ± 1.5		MURTHY	89	Cesium, no <i>B</i> field
– 50 ± 110		LAMOREAUX	87	^{199}Hg
190 ± 340	90	SANDARS	75	Thallium
70 ± 220	90	PLAYER	70	Xenon
< 300	90	WEISSKOPF	68	Cesium

13 ABDULLAH 90, COMMINS 94, and REGAN 02 use the relativistic enhancement of a valence electron's electric dipole moment in a high-Z atom.

e^- MEAN LIFE / BRANCHING FRACTION

A test of charge conservation. See the "Note on Testing Charge Conservation and the Pauli Exclusion Principle" following this section in our 1992 edition (Physical Review **D45**, 1 June, Part II (1992), p. VI.10).

Most of these experiments are one of three kinds: Attempts to observe (a) the 255.5 keV gamma ray produced in $e^- \rightarrow \nu_e \gamma$, (b) the (K) shell x ray produced when an electron decays without additional energy deposit, e.g., $e^- \rightarrow \nu_e \bar{\nu}_e \nu_e$ ("disappearance" experiments), and (c) nuclear de-excitation gamma rays after the electron disappears from an atomic shell and the nucleus is left in an excited state. The last can include both weak boson and photon mediating processes. We use the best $e^- \rightarrow \nu_e \gamma$ limit for the Summary Tables.

Note that we use the mean life rather than the half life, which is often reported.

$e \rightarrow \nu_e \gamma$ and astrophysical limits

<u>VALUE</u> (yr)	<u>CL%</u>	<u>DOCUMENT ID</u>	<u>TECN</u>	<u>COMMENT</u>
$>4.6 \times 10^{26}$	90	BACK 02	BORX	$e^- \rightarrow \nu \gamma$

• • • We do not use the following data for averages, fits, limits, etc. • • •

$>3.4 \times 10^{26}$	68	BELLI	00B DAMA	$e^- \rightarrow \nu\gamma$, liquid Xe
$>3.7 \times 10^{25}$	68	AHARONOV	95B CNTR	$e^- \rightarrow \nu\gamma$
$>2.35 \times 10^{25}$	68	BALYSH	93 CNTR	$e^- \rightarrow \nu\gamma$, ^{76}Ge detector
$>1.5 \times 10^{25}$	68	AVIGNONE	86 CNTR	$e^- \rightarrow \nu\gamma$
$>1 \times 10^{39}$	¹⁴ ORITO		85 ASTR	Astrophysical argument
$>3 \times 10^{23}$	68	BELLOTTI	83B CNTR	$e^- \rightarrow \nu\gamma$

¹⁴ ORITO 85 assumes that electromagnetic forces extend out to large enough distances and that the age of our galaxy is 10^{10} years.

Disappearance and nuclear-de-excitation experiments

VALUE (yr)	CL%	DOCUMENT ID	TECN	COMMENT
$>6.4 \times 10^{24}$	68	¹⁵ BELLI	99B DAMA	De-excitation of ^{129}Xe
• • • We do not use the following data for averages, fits, limits, etc. • • •				
$>4.2 \times 10^{24}$	68	BELLI	99 DAMA	Iodine L-shell disappearance
$>2.4 \times 10^{23}$	90	¹⁶ BELLI	99D DAMA	De-excitation of ^{127}I (in NaI)
$>4.3 \times 10^{23}$	68	AHARONOV	95B CNTR	Ge K-shell disappearance
$>2.7 \times 10^{23}$	68	REUSSER	91 CNTR	Ge K-shell disappearance
$>2 \times 10^{22}$	68	BELLOTTI	83B CNTR	Ge K-shell disappearance

¹⁵ BELLI 99B limit on charge nonconserving e^- capture involving excitation of the 236.1 keV nuclear state of ^{129}Xe ; the 90% CL limit is 3.7×10^{24} yr. Less stringent limits for other states are also given.

¹⁶ BELLI 99D limit on charge nonconserving e^- capture involving excitation of the 57.6 keV nuclear state of ^{127}I . Less stringent limits for the other states and for the state of ^{23}Na are also given.

e REFERENCES

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REGAN	02	PRL 88 071805	B.C. Regan <i>et al.</i>	
BELLI	00B	PR D61 117301	P. Belli <i>et al.</i>	(DAMA Collab.)
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AHARONOV	95B	PR D52 3785	Y. Aharonov <i>et al.</i>	(SCUC, PNL, ZAGR+)
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ORITO	85	PRL 54 2457	S. Orito, M. Yoshimura	(TOKY, KEK)
CHU	84	PRL 52 1689	S. Chu, A.P. Mills, J.L. Hall	(BELL, NBS, COLO)
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PLAYER	70	JPB 3 1620	M.A. Player, P.G.H. Sandars	(OXF)
WEISSKOPF	68	PRL 21 1645	M.C. Weisskopf <i>et al.</i>	(BRAN)